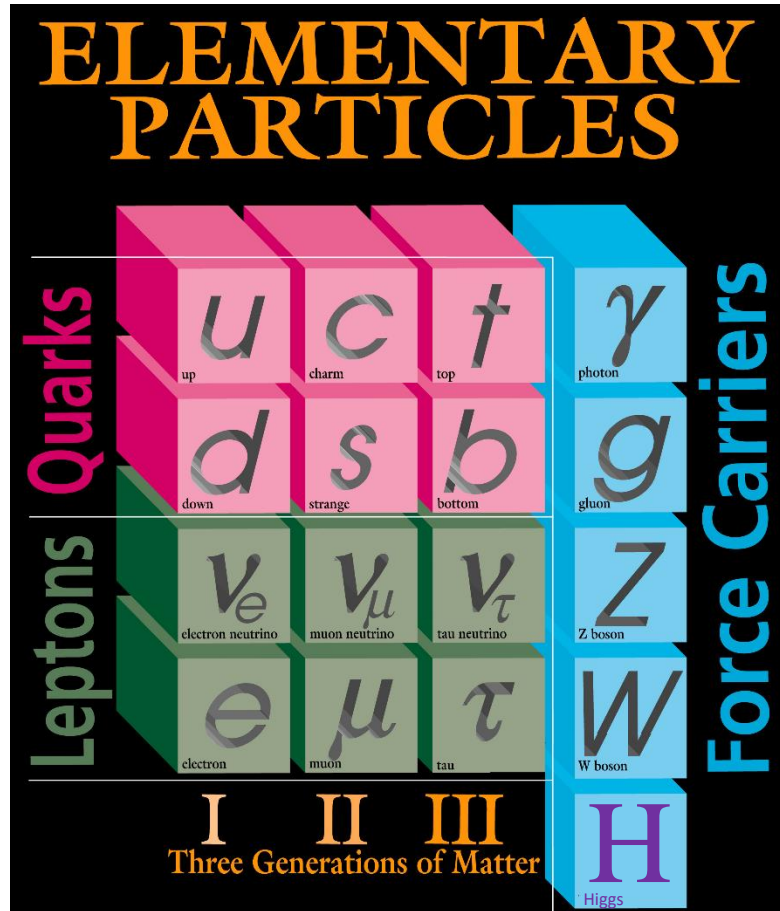


NOvA Experiment

Neutrino Activity



Building Blocks of the universe



Standard Model of Particle Physics

Force carrying bosons

Three generations of fermions

Quarks

3 positively charged quarks

3 negatively charged quarks

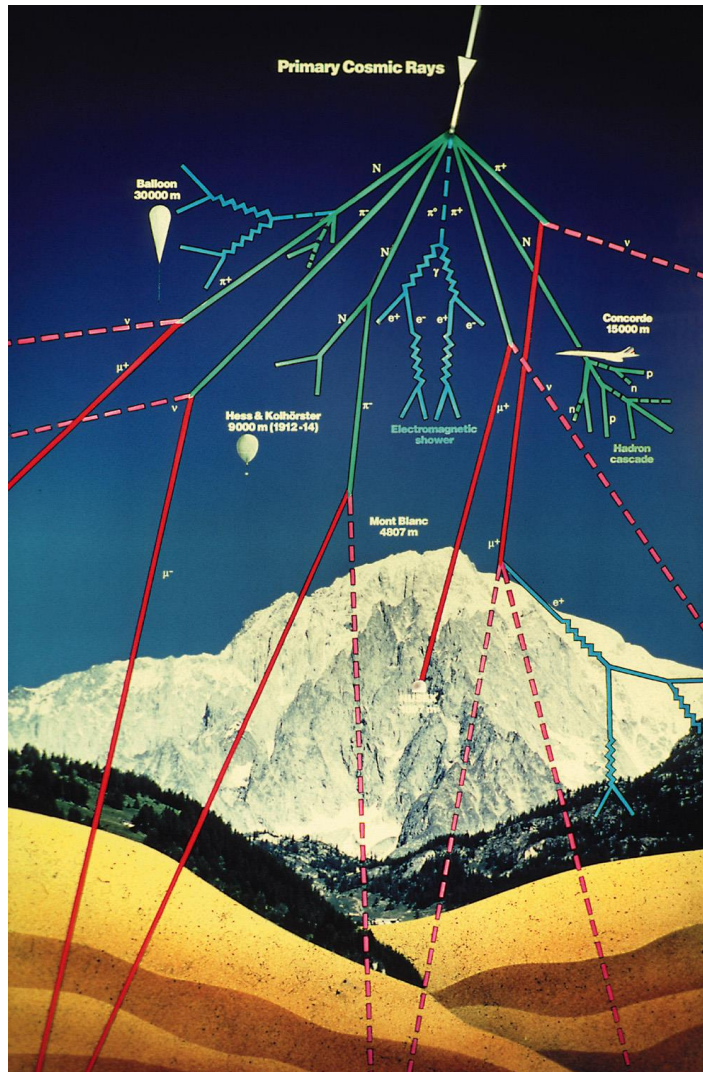
Make up protons, neutrons, etc

Leptons

3 charged leptons

3 neutral leptons

Heavy cousins of the electron



Heavier versions of the electron

Muon – μ

200 times the electron mass

Lifetime of $2 \mu\text{s}$

Interact with us as cosmic rays

$\Theta(1)$ through hand each second

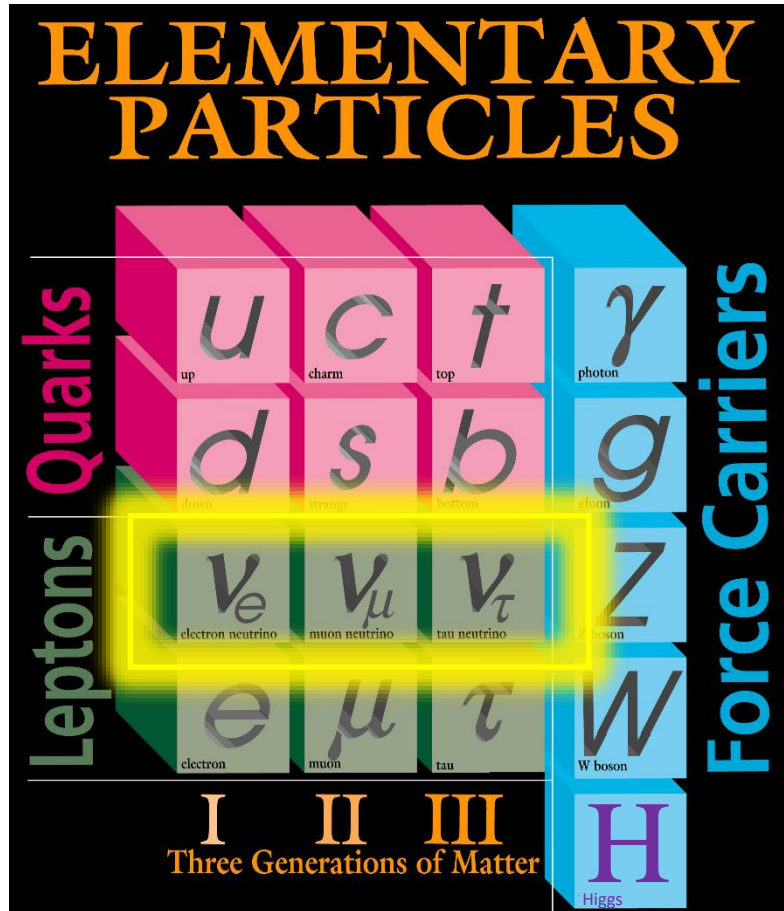
$\Theta(10\%)$ daily radiation exposure

Tau – τ

3,500 times the electron mass

Lifetime of 0.3 ps

What about the neutrinos



Neutrinos are probably the least understood Standard Model particle

We've known about them for ~100 years

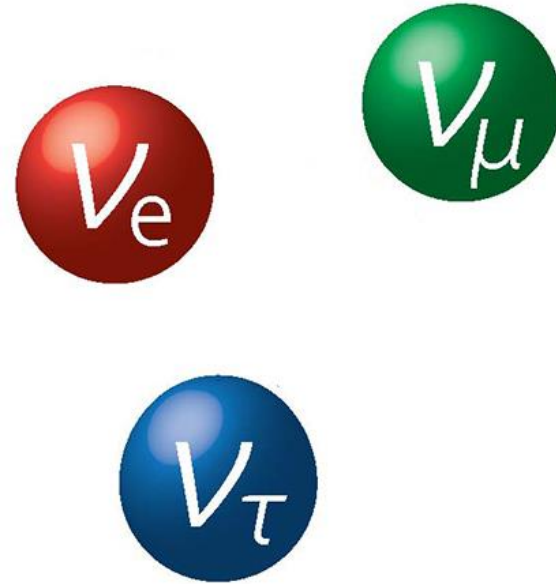
2nd most abundant particle in universe

$\Theta(100 \text{ million})$ per cubic meter

100 billion solar neutrinos pass

through your finger nail each second

What about the neutrinos



How they're different

No electric or color charge

No EM or strong interactions

Extremely tiny mass

At least a million times less massive than an electron

Only interact through weak force and gravity

Gravity is so weak

Only observed in cosmology

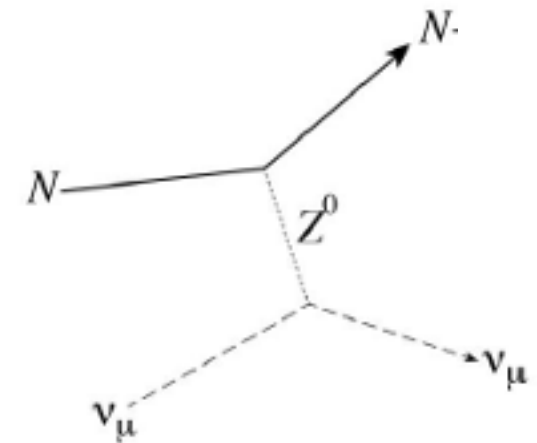
Completely ignorable at particle level

Weak interaction important at particle level

Mediated by W & Z bosons with mass

$M_{Z,W}$ typically greater than ν energy

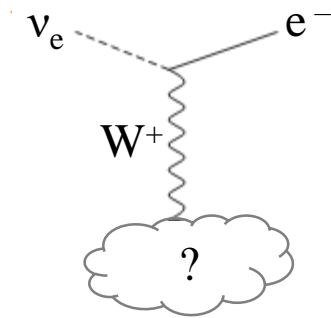
Think Heisenberg Uncertainty → Rare chance to create mass energy → Rare chance to interact



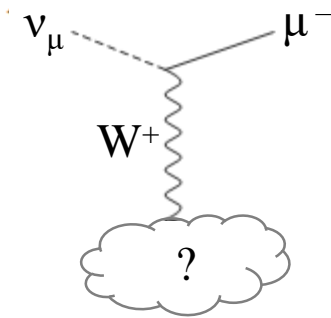
Neutrino Flavor

Identify neutrino flavor through charged-current (CC) interaction

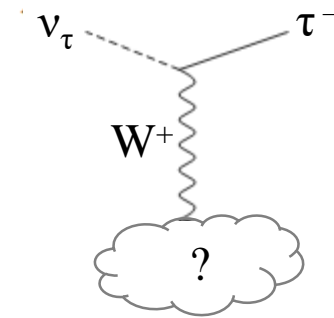
Electron Neutrino



Muon Neutrino

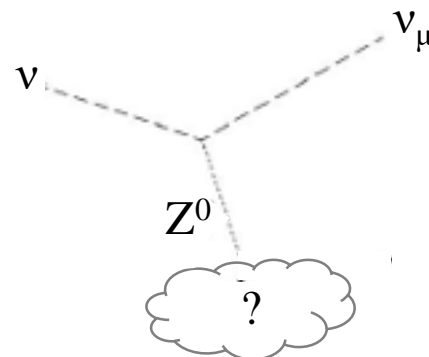


Tau Neutrino




Neutral Current (NC) Interaction

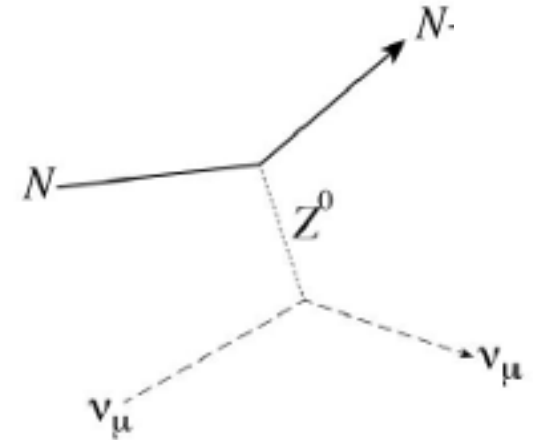
All neutrino flavors look the same



Neutrinos Rarely Interact

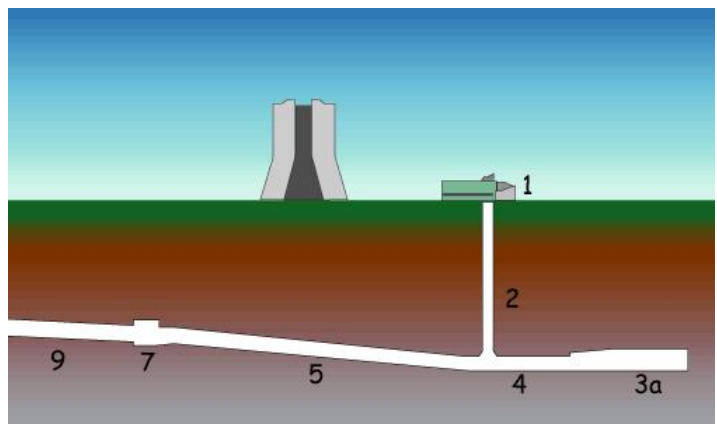
Shoot a neutrino through a infinite long slab of lead
It most likely won't interact until it has spent
over 1 year traveling through the material

ν 
 ν at typical energies

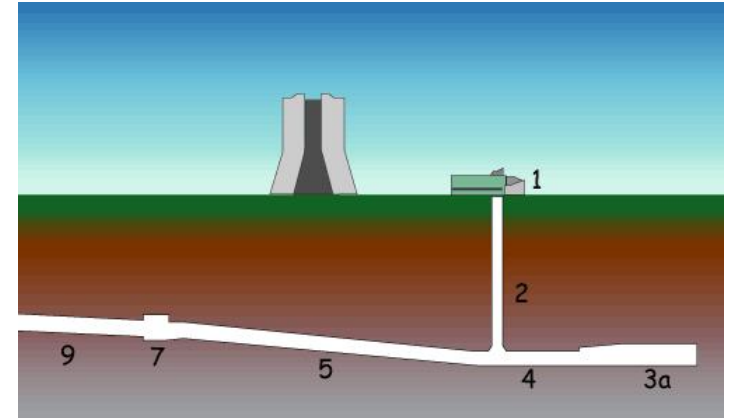
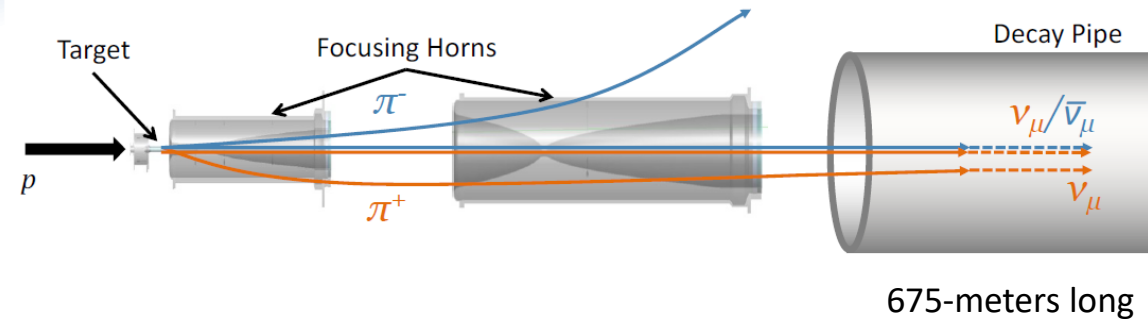


There is some probability that a neutrino will interact at any point
The more neutrinos that pass through a point, the more chances that one will interact at that point

NOvA Experiment

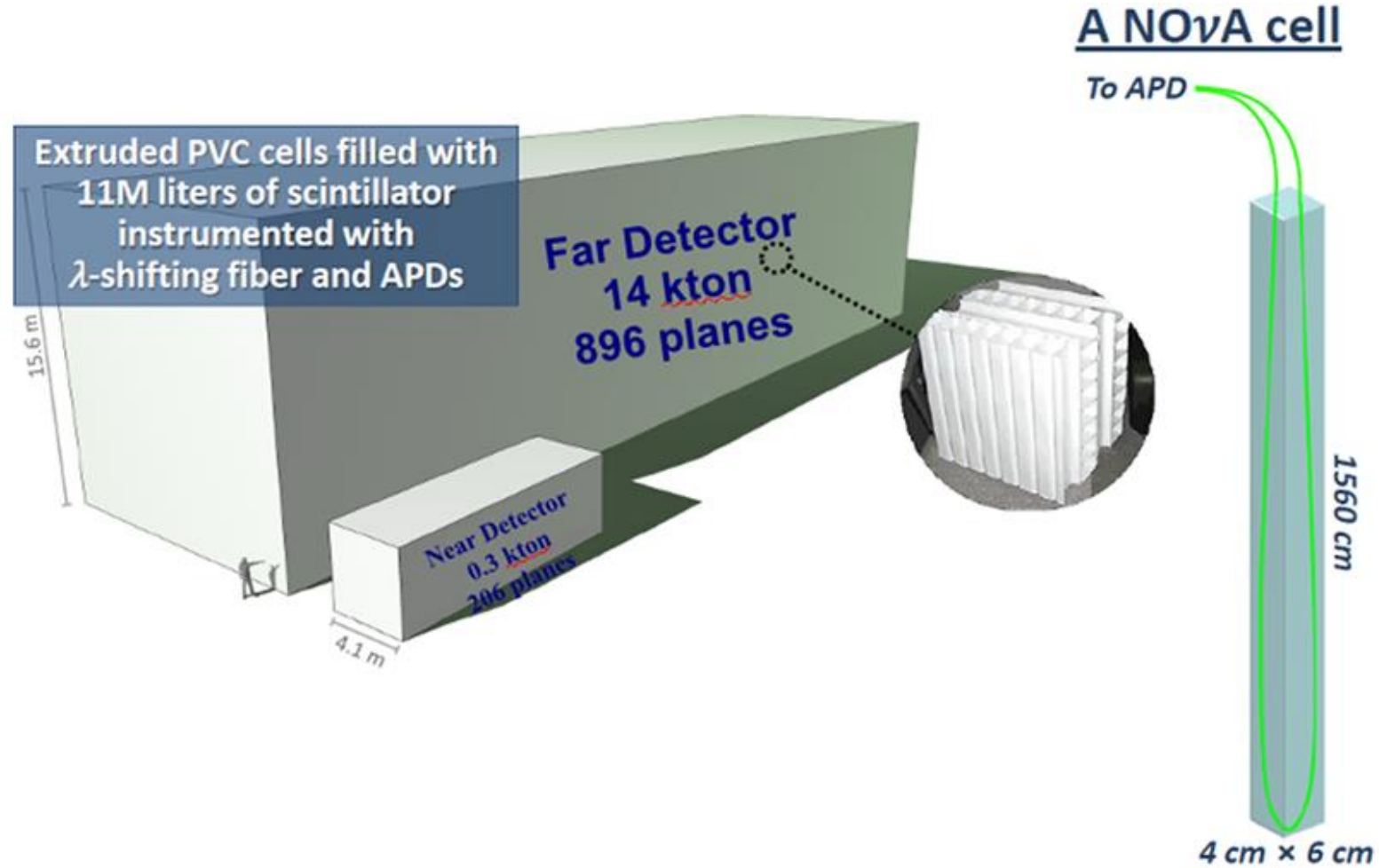
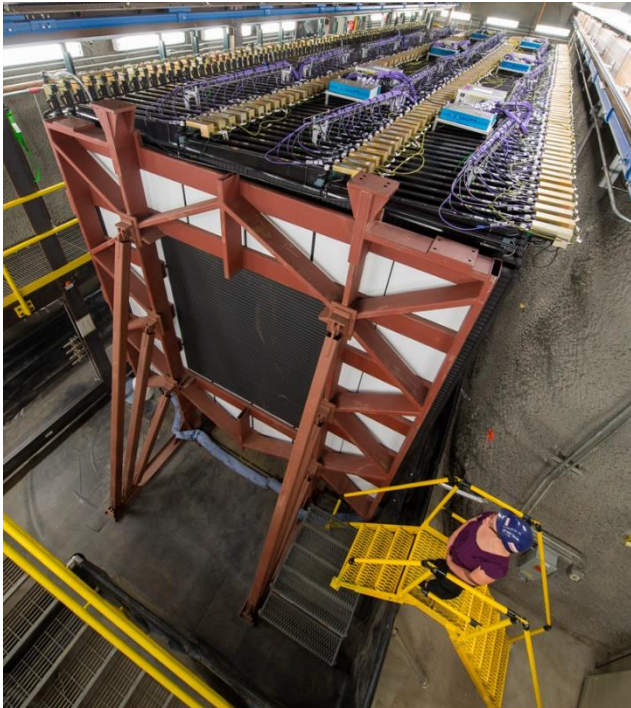
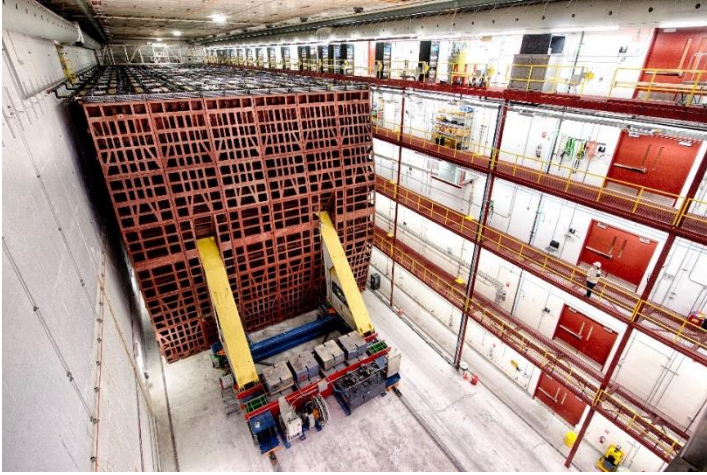


NOvA Neutrino Beam

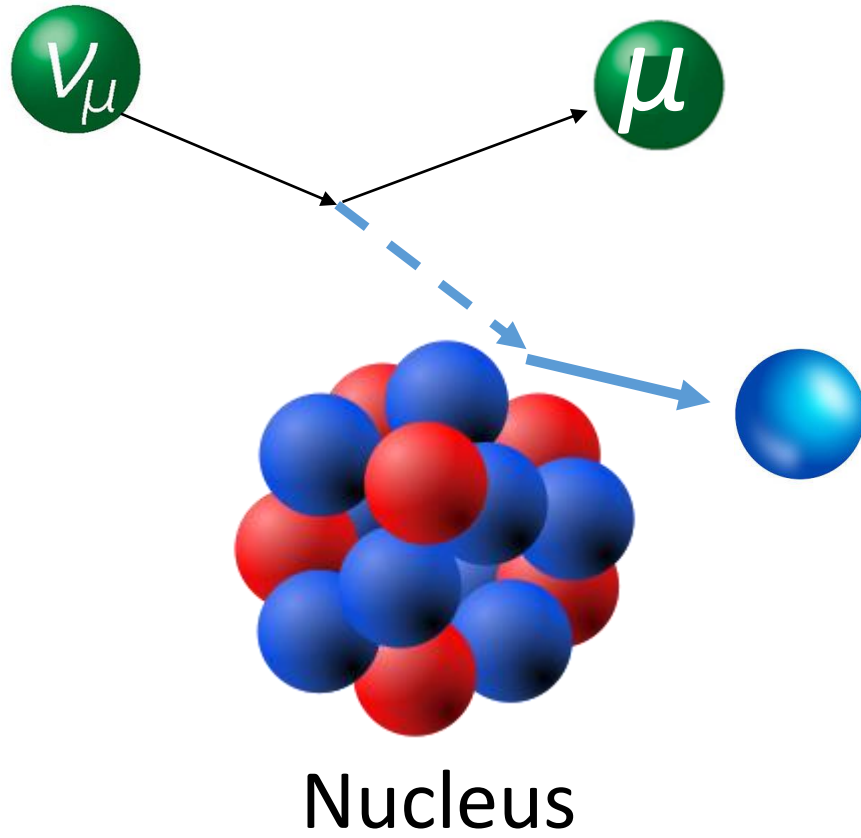


Number of neutrinos produced is proportional to the number of protons on target

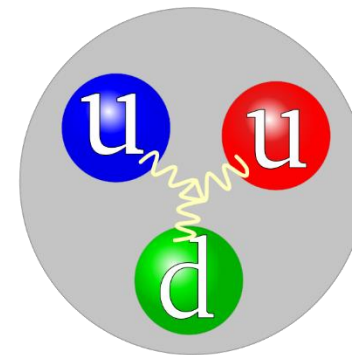
NOvA Detectors



Neutrino Interactions



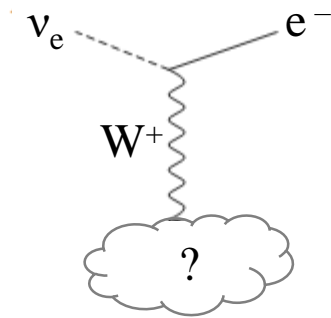
Proton



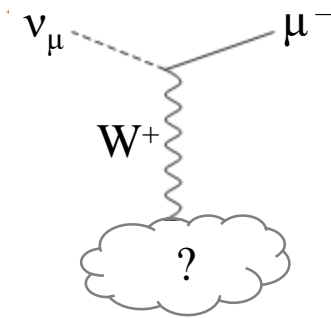
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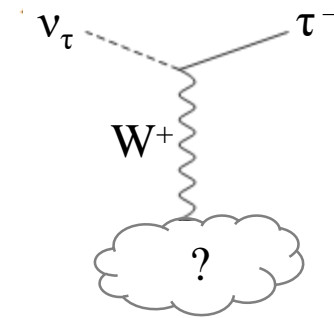
Electron Neutrino



Muon Neutrino

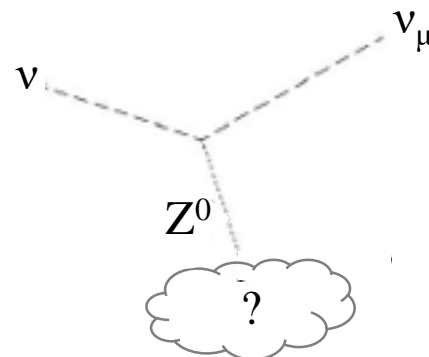


Tau Neutrino

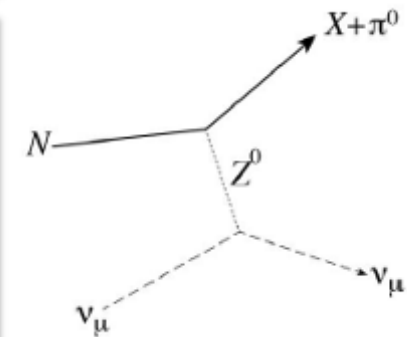
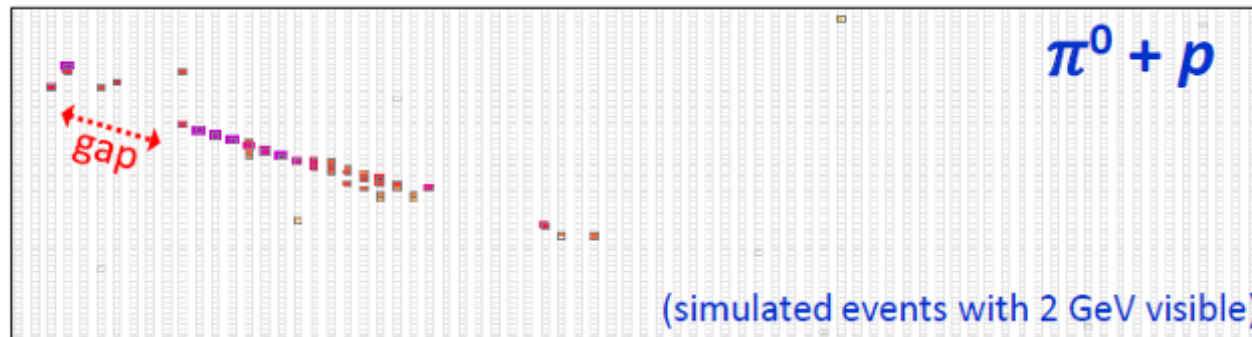
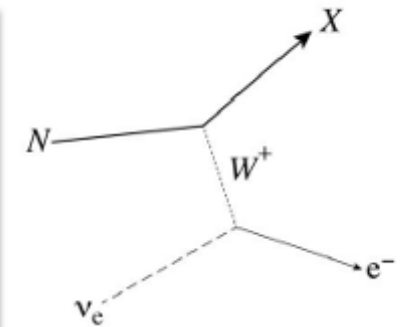
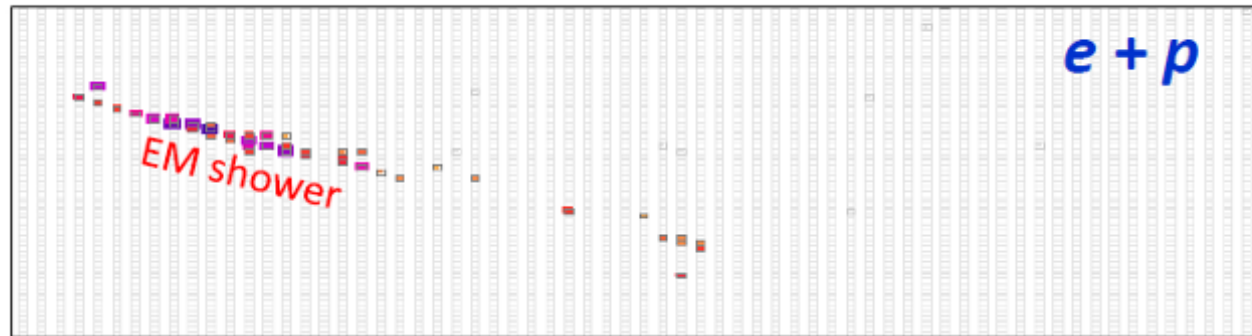
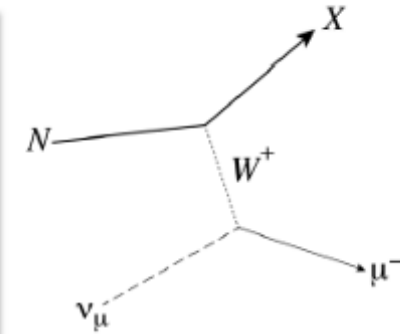
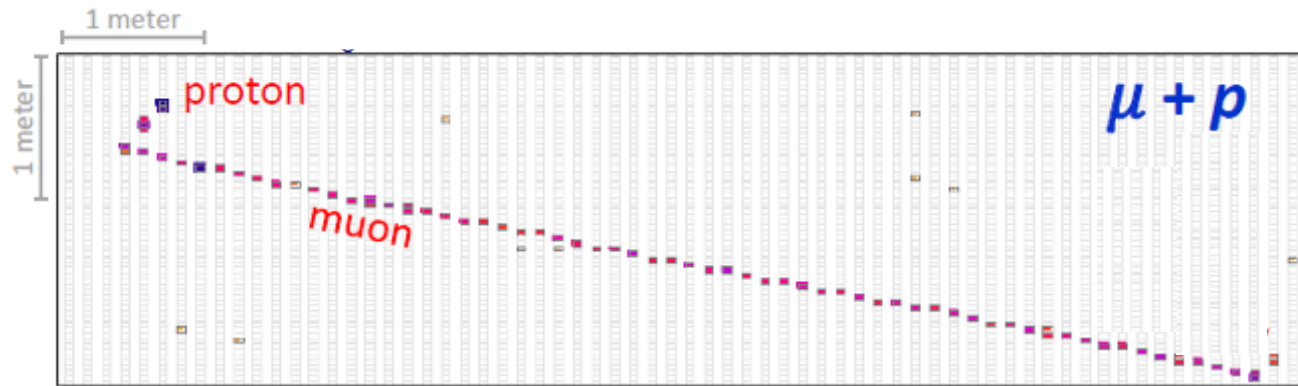


Neutral Current (NC) Interaction

All neutrino flavors look the same



Neutrino Interactions



Let's look at FD events

- Students form groups of two and look at FD event displays
- Note for the same exposure of protons-on-target for the NOvA FD data used for this study there are:
 - 6 NuMu-CC events – only the numu flavor can cause these events
 - 40 NC events – any neutrino flavor can cause these events

Relative to a specific selection criteria applied to data sample
- By looking at the images can each group come up with some **quantifiable** criterion (or criteria) that could be used to distinguish NuMC-CC events from NC events
- Discuss groups' thoughts on what **quantifiable** criterion could be applied

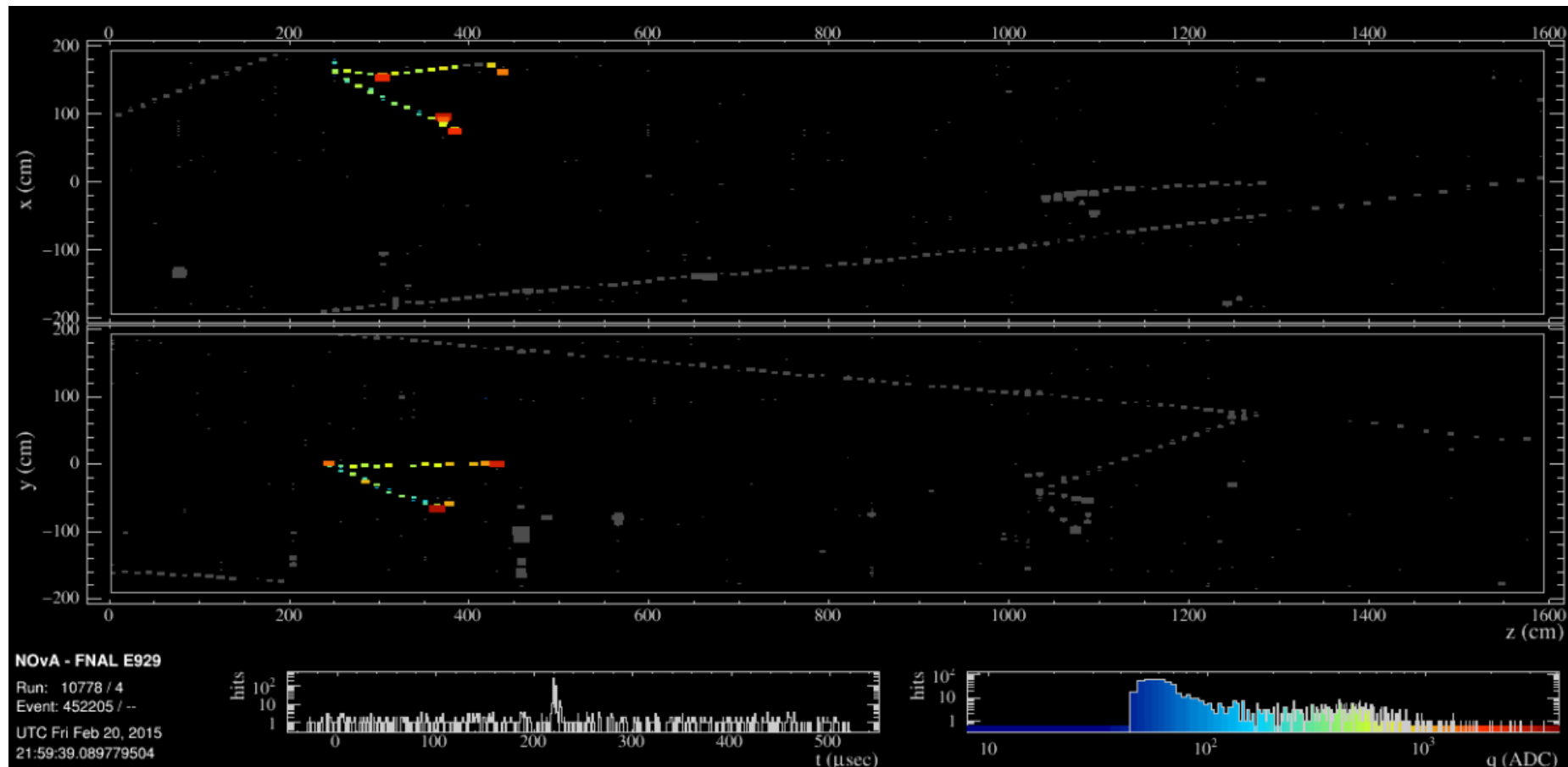
About the event displays

There are two views

- One in which all of the cells are vertical

- One in which all of the cells are horizontal

Color indicates energy that is deposited in the cell by the particle



Work as a group

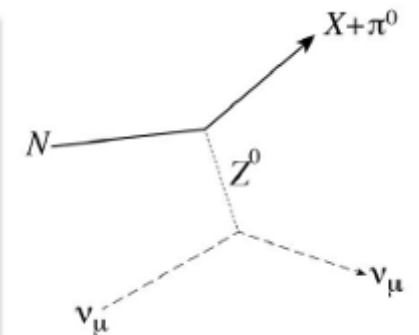
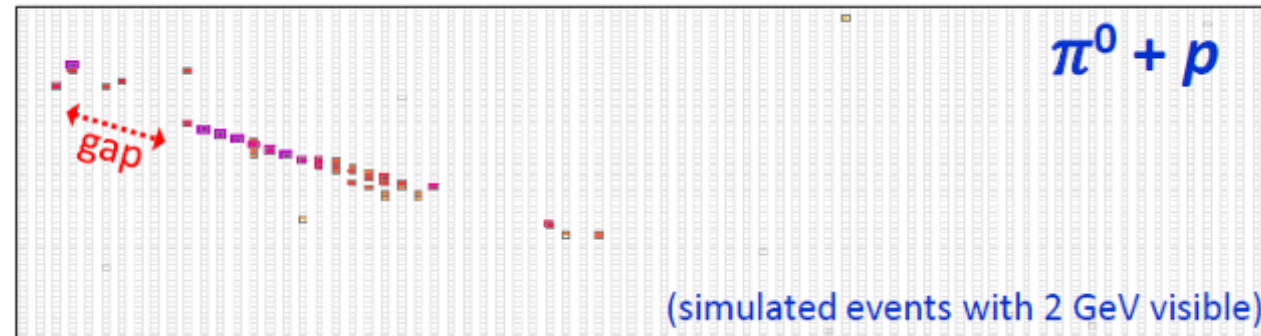
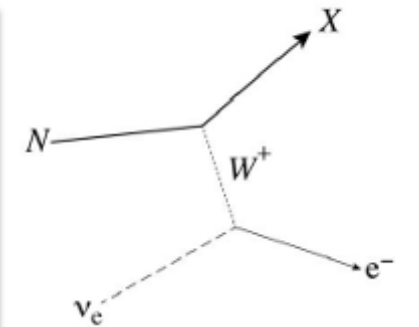
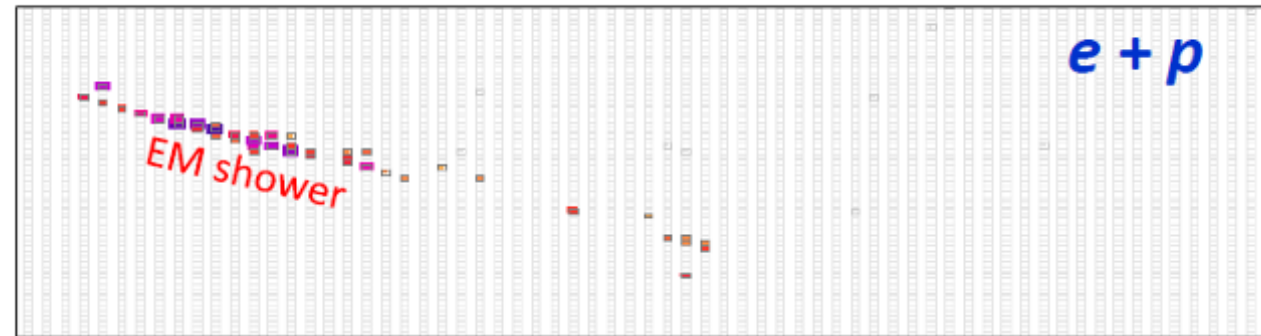
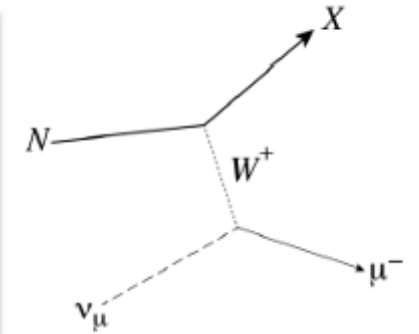
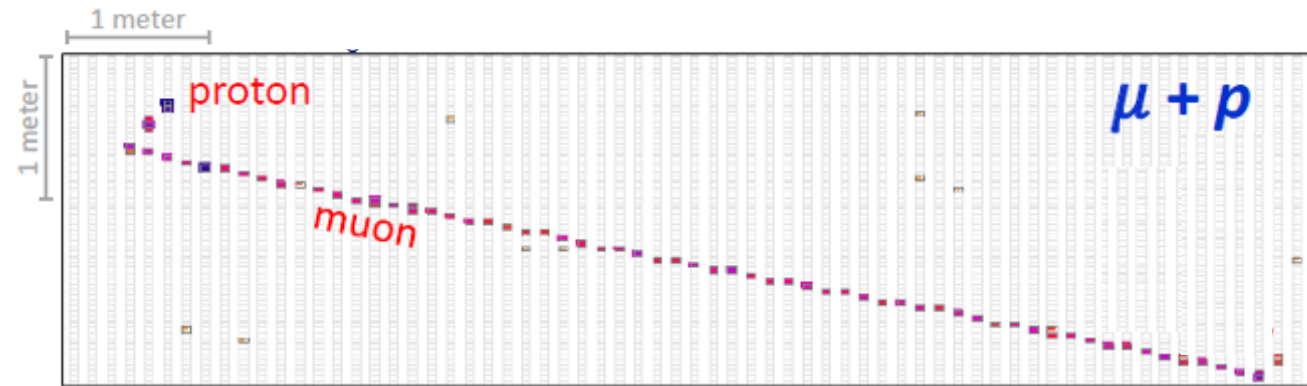
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Neutrino Interactions

NuMu's are identifiable through presence of a muon that make a long straight track



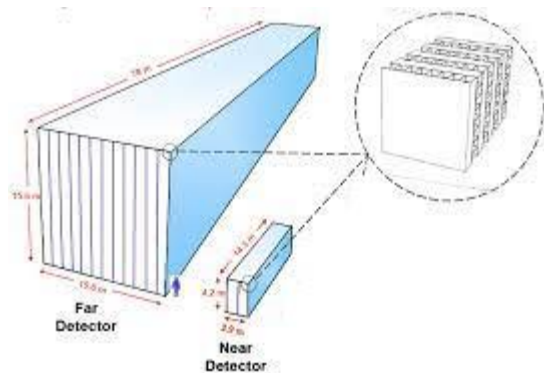
Spreadsheet of ND events

The ND events are collected using a data sample that corresponds to $1/50^{\text{th}}$ the protons-on-target as the FD.

Question for students: Think about what you might expect about a spread sheet with ND events, that corresponds to $1/50^{\text{th}}$ of the beam exposure as the FD. Then open the spread sheet, and discuss if anything meets your expectations or not?

Spreadsheet of ND events

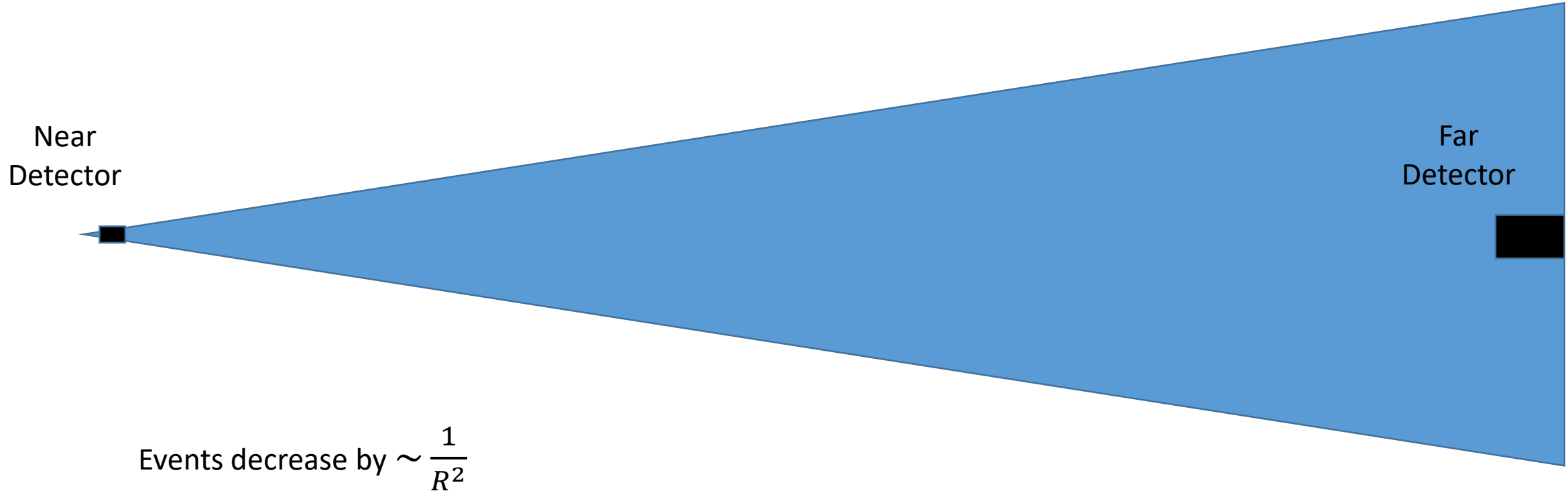
Question for students: Why are there so many events in the ND file compared to the FD even though the ND files is a much smaller amount of protons-on-target?



Near Detector

The near detector is much closer to the neutrino source than the far detector
→ greater flux of neutrinos passing through the near detector

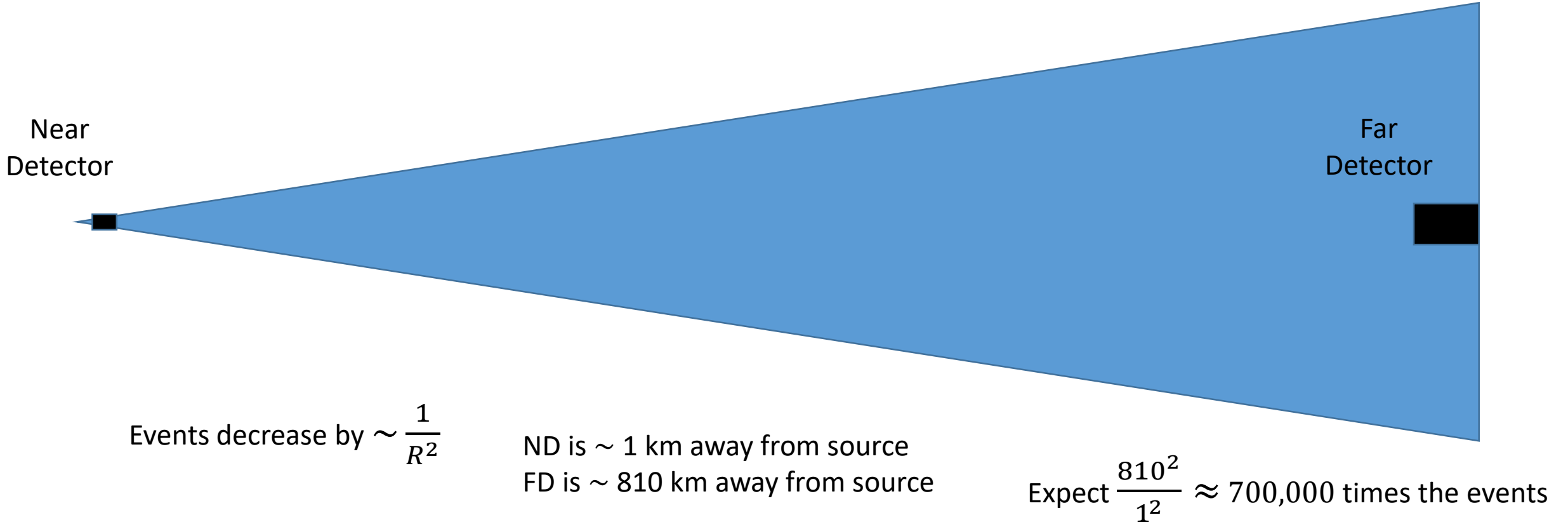
Think about a flash light shining on a wall when you are close and far away.



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Order of magnitude estimates

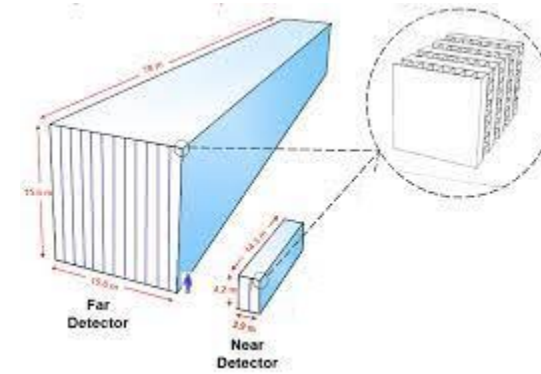
Near Detector

Expect 700,000 times the events

But the ND is much less massive than the FD

Neutrinos interact with proportionally less nucleons

Particles produced in neutrino interactions are more likely to escape smaller ND. Only look at events that are fully contained



Near Detector

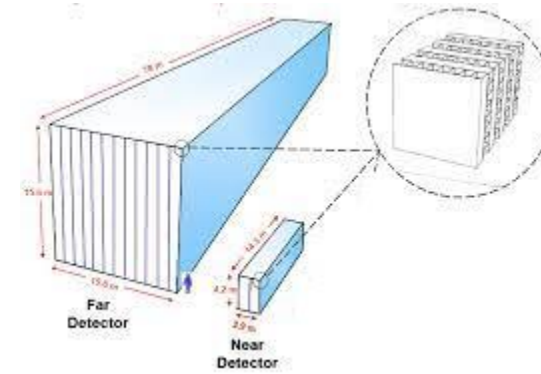
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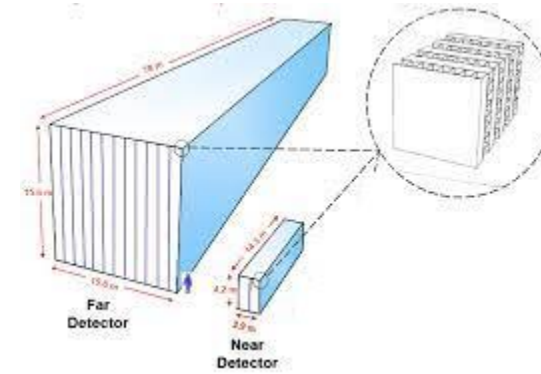
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Roughly expect $700,000/170 \approx 4000$ times more events in ND than FD for same POT



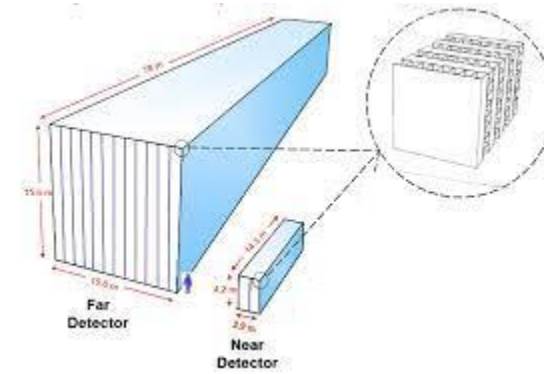
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The ND has $1/50^{\text{th}}$ the protons-on-target, so we roughly expect $4000/50 \approx 80$ more events in ND than FD

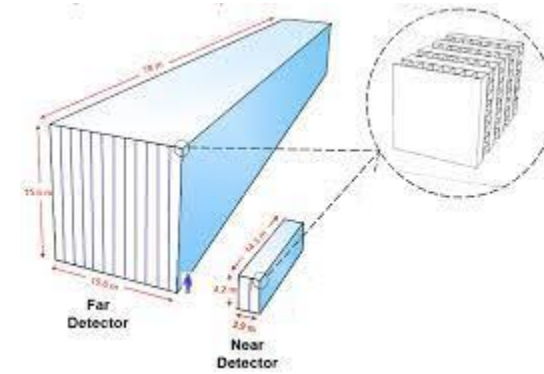
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Does our ND sample match our order of magnitude estimate?

How much of the ND events are CC vs NC?

- Do some coding to identify CC vs NC

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- Do results match your expectations?

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Neutrinos change flavor as they travel!

We call this neutrino oscillation, since they will eventually change back to their original flavor